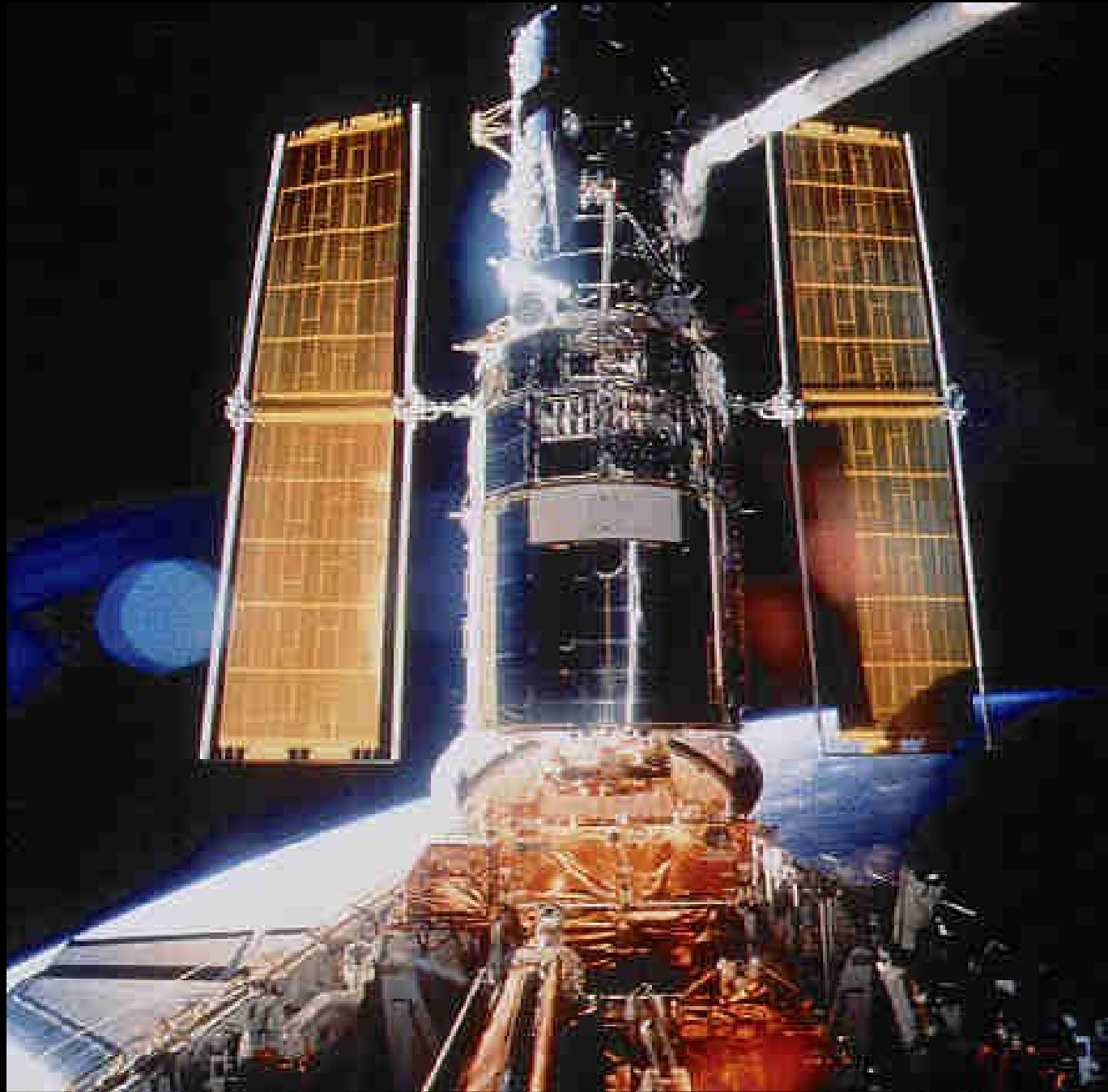
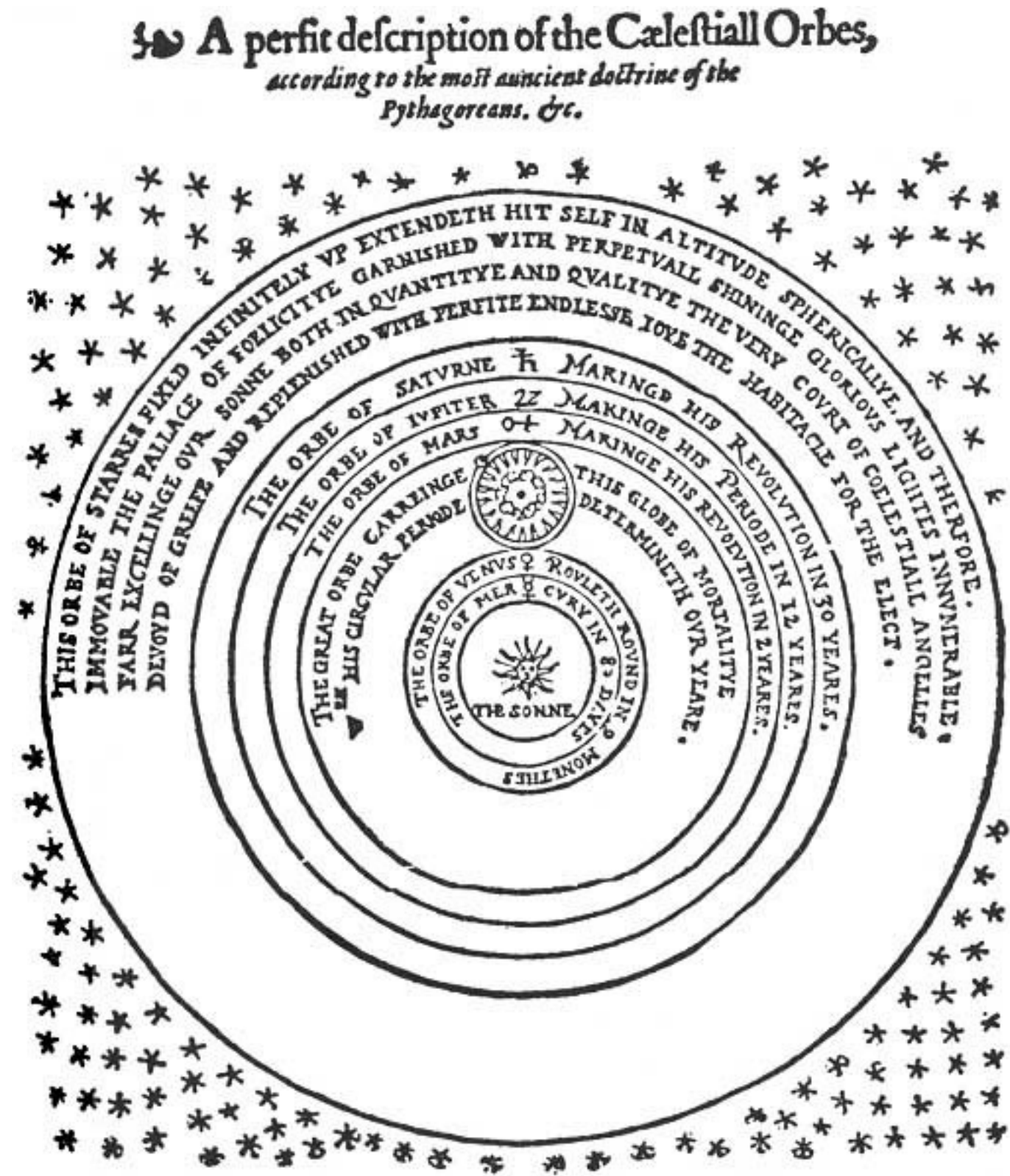


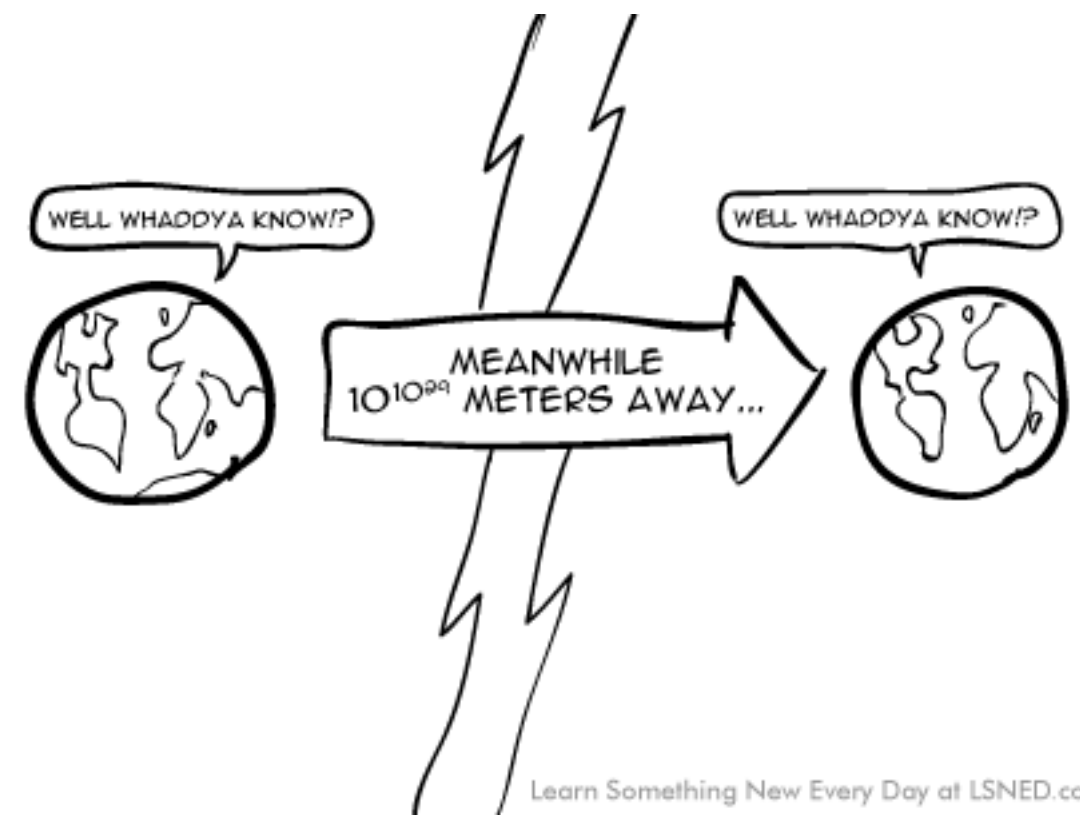
# Cosmology



# Newton's model

- According to Newton, the universe was:
- Infinite (in time and space)
- Uniform (same everywhere)
- Static (not moving)

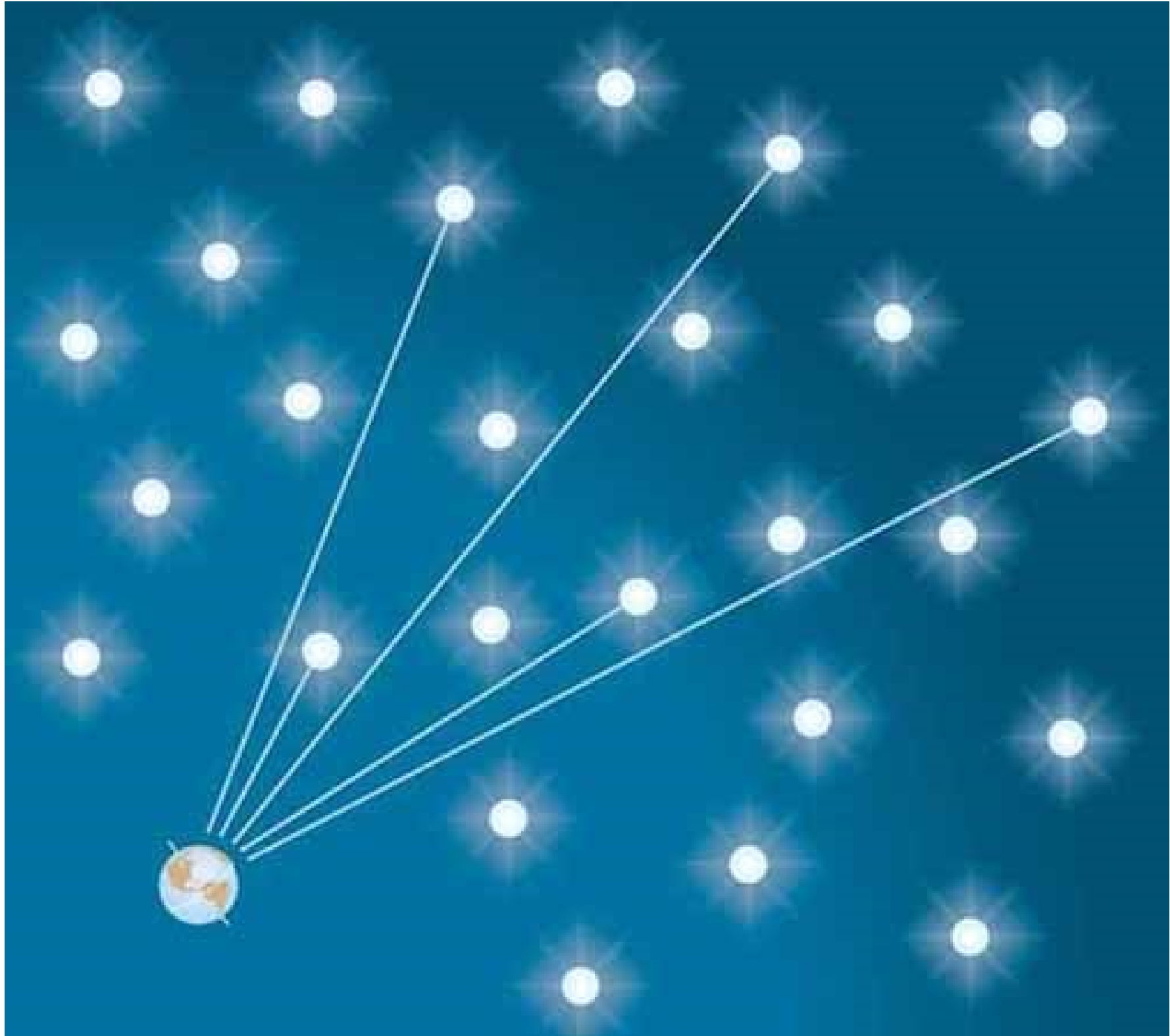




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# Why is the night sky dark?



# Olber's paradox

- If the universe is infinite in size, there must be an infinite number of stars
- Every line of sight should terminate in the surface of a star. Sunglasses anyone?
- We can resolve this in two ways:
- The universe has a finite age, therefore light from distant stars hasn't reached us yet
- The universe is expanding, stretching radiation into the infrared.

# Hubble

- Using the biggest telescope in the world, Edwin Hubble discovered some “nebulae” were actually “island universes” or what we now call galaxies



# Hubble activity

Use the data on my website to  
graph recession velocity vs.  
distance

Use the slope of the graph to  
find the age of the universe





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# Three main types of galaxies:

- Elliptical
- Spiral
- Irregular



# Elliptical Galaxies

- These have a round or slightly elliptical shape
- They tend to be red in color



# Spiral Galaxies



- Their rotation helps us understand their spiral arms and mass distribution
- They can be classified by their “tightness” and bars



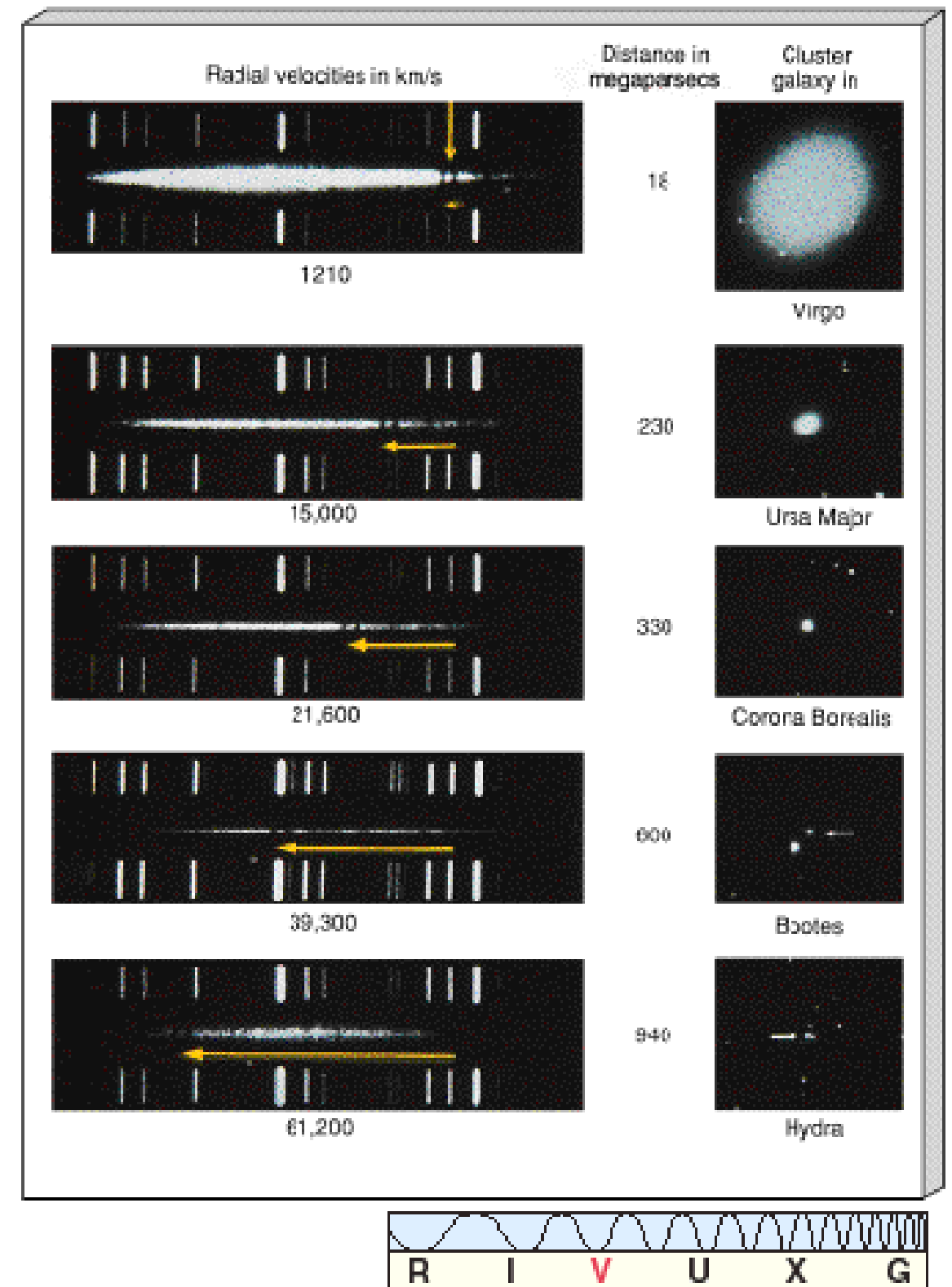
# Irregular Galaxies:

- These form when two galaxies collide



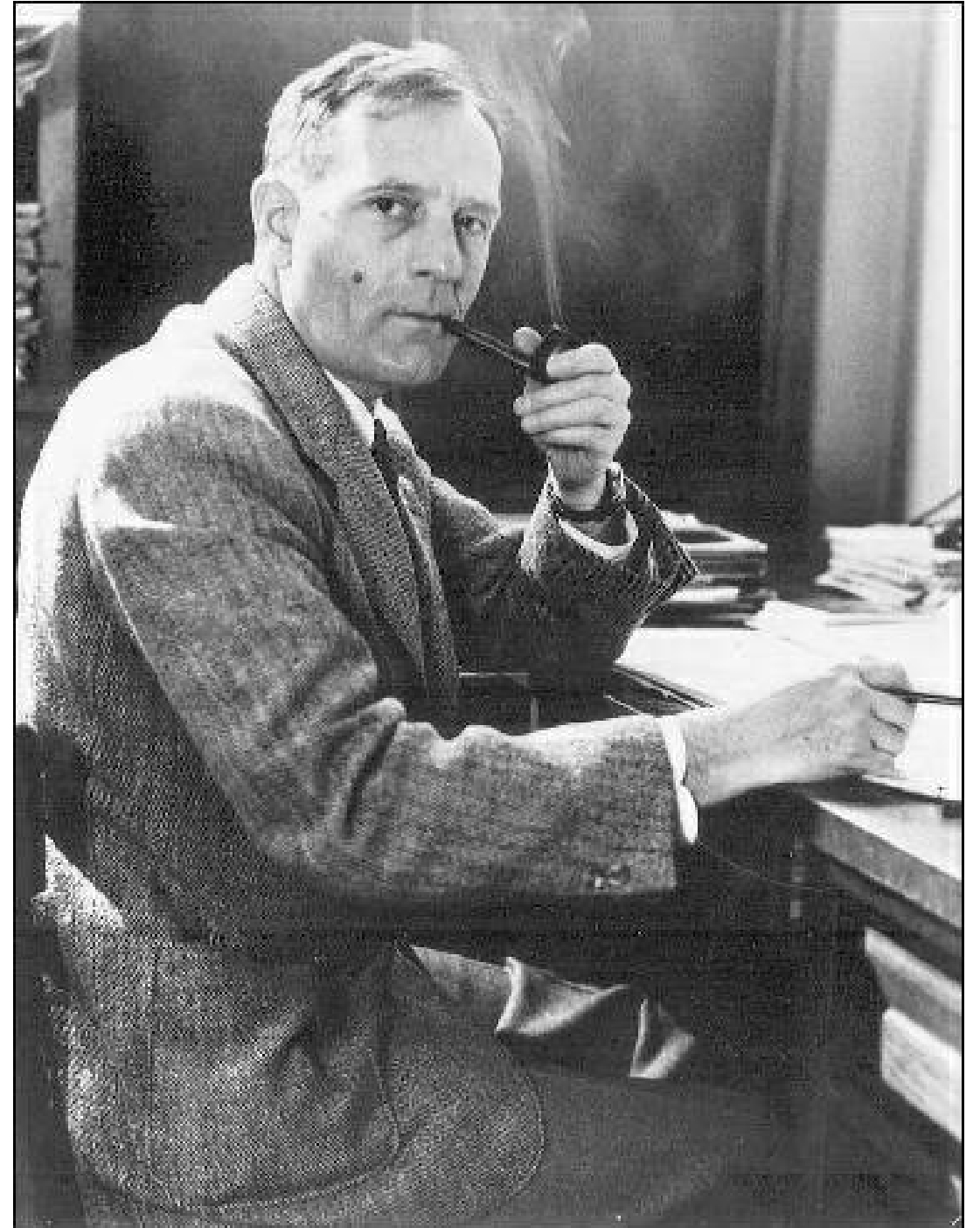
# The Doppler Effect

- All distant galaxies were observed to be moving away from us
- The further the galaxy, the larger the redshift



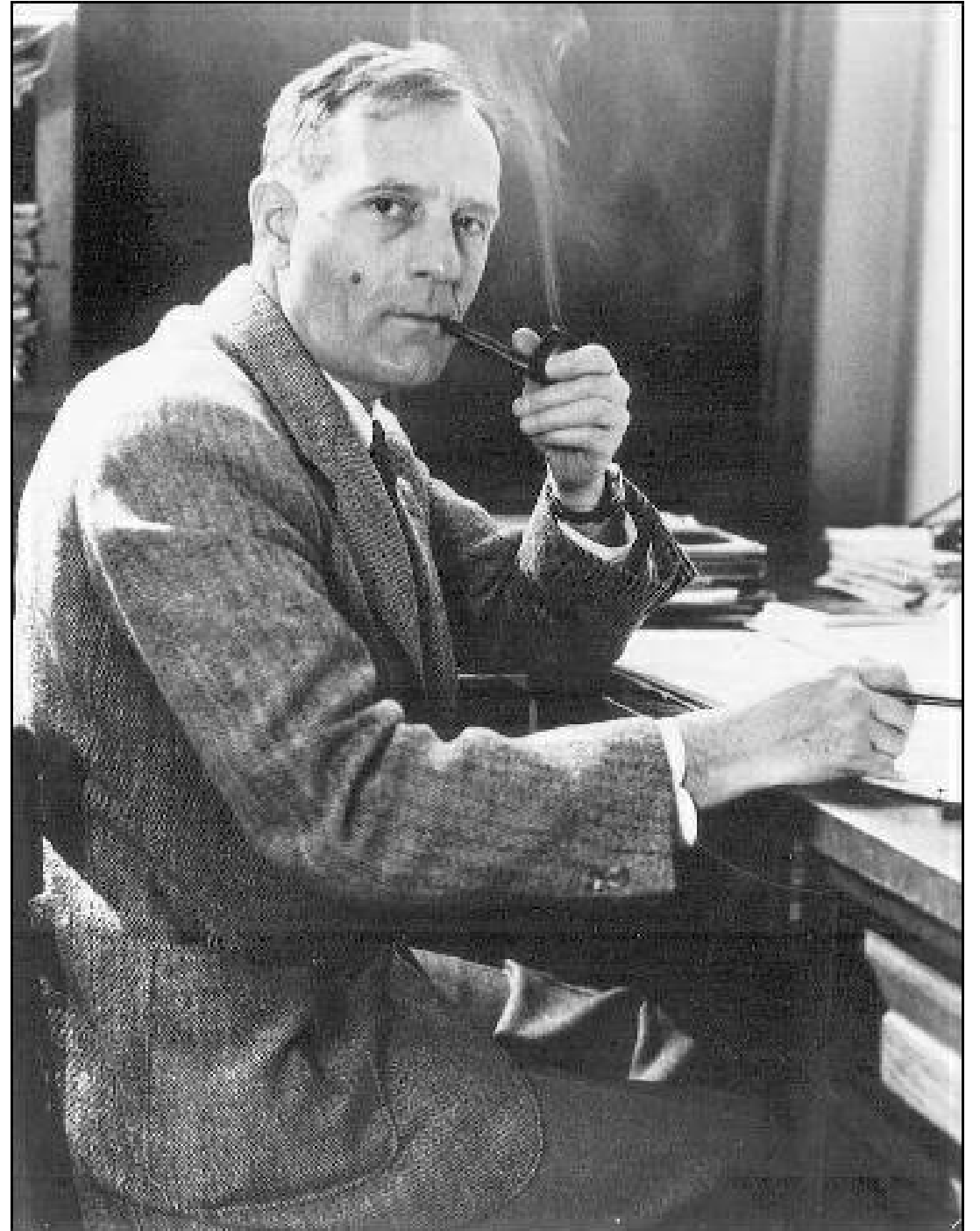
# Inflation!

- Hubble then realized the universe must be expanding to account for this “cosmological redshift”



# Hubble's law

- Use the data on my website to graph distance vs speed





- 
- A deep-field astronomical image showing a vast field of galaxies and distant stars against a black background. The galaxies are of various shapes and sizes, some appearing as bright, elongated structures, while others are smaller, distant points of light. The colors range from bright yellow and white to deep reds and blues, indicating different temperatures and compositions. The overall scene is a dense, colorful mosaic of cosmic objects.
- Astronomers now believe the universe will not only continue but accelerate!



# Redshift $z$

Ex: Find the recessional velocity of a galaxy if a 440 nm spectral line is redshifted by 23 nm.  
Redshift  $z$  is given by:

$$z = \frac{\Delta\lambda}{\lambda} = \frac{v}{c}$$

$$z = \frac{23}{440} = 0.0523$$

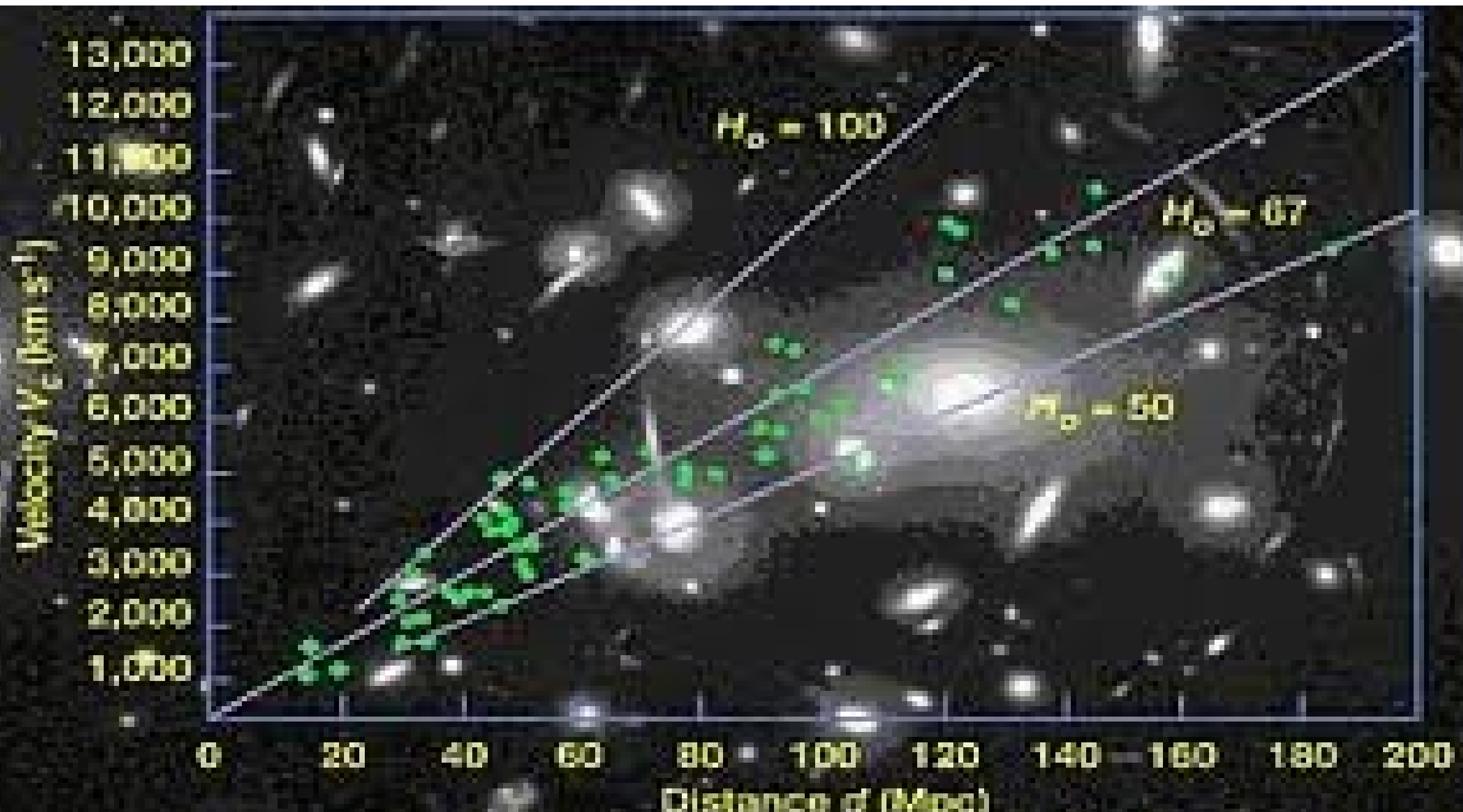
$$v = 0.0523c = 1.6 \times 10^7 \text{ m} \cdot \text{s}^{-1}$$

# Hubble's Law

- If we know the redshift of receding galaxies and how far they are away, we can calculate the...
- Hubble constant!!!
- Hubble's law tells us the further a galaxy is, the faster it is receding from us:
- $v = H_0 d$

$$H_0 = 71 \frac{\text{km} / \text{s}}{\text{Mpc}} = 2.3 \times 10^{-18} \text{ s}^{-1}$$

$$t_H = \frac{1}{2.3 \times 10^{-18} \text{ s}^{-1}} = 13.8 \times 10^9 \text{ years}$$



# Redshift example

Ex: What wavelength would we expect to observe for the H-alpha Balmer emission line from a galaxy 11 billion light years away (use  $H=71\text{km}\cdot\text{s}^{-1}\cdot\text{Mpc}^{-1}$ )?

$$v = H_0 d \quad \frac{\Delta\lambda}{\lambda} = \frac{71000\text{m}\cdot\text{s}^{-1} \cdot \text{Mpc} \cdot 3374\text{Mpc}}{3 \cdot 10^8\text{m}\cdot\text{s}^{-1}}$$

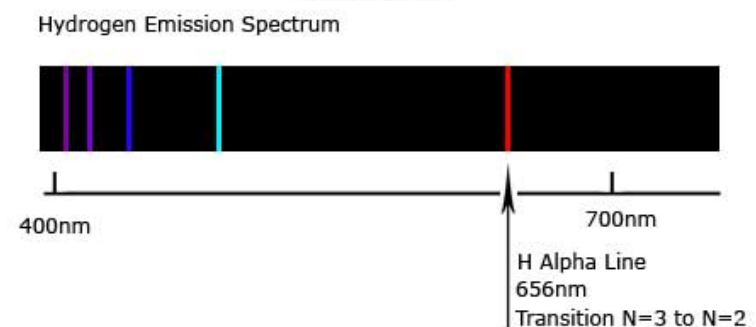
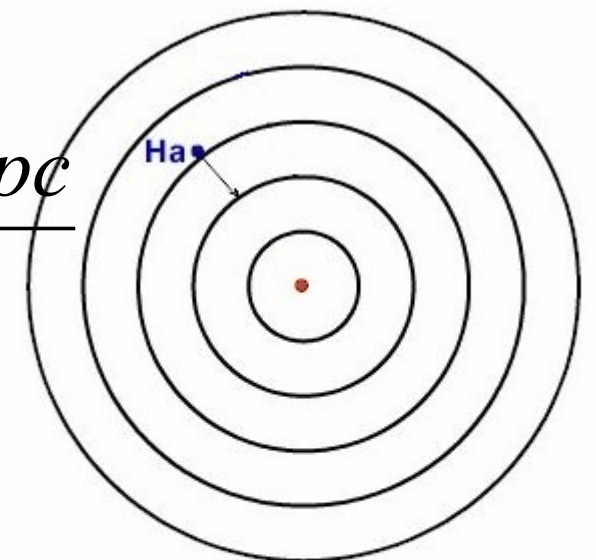
$$z = \frac{\Delta\lambda}{\lambda} = \frac{v}{c}$$

$$\Delta\lambda = 0.798 \cdot 656.3\text{nm}$$

$$\Delta\lambda = 524\text{nm}$$

$$\frac{\Delta\lambda}{\lambda} = \frac{H_0 d}{c}$$

$$\lambda = 524 + 656 = 1180\text{nm}$$



# Cosmic Scale Factor R

As the universe expands, scale factor R increases.

This provides an alternate explanation to doppler.

Ex: The peak wavelength of the CMB radiation at present is about 1 mm. In the past there was a time when the peak wavelength corresponded to blue light (400 nm). Estimate the scale factor ratio.

$$z = \frac{\Delta\lambda}{\lambda_0} = \frac{R}{R_0} - 1$$



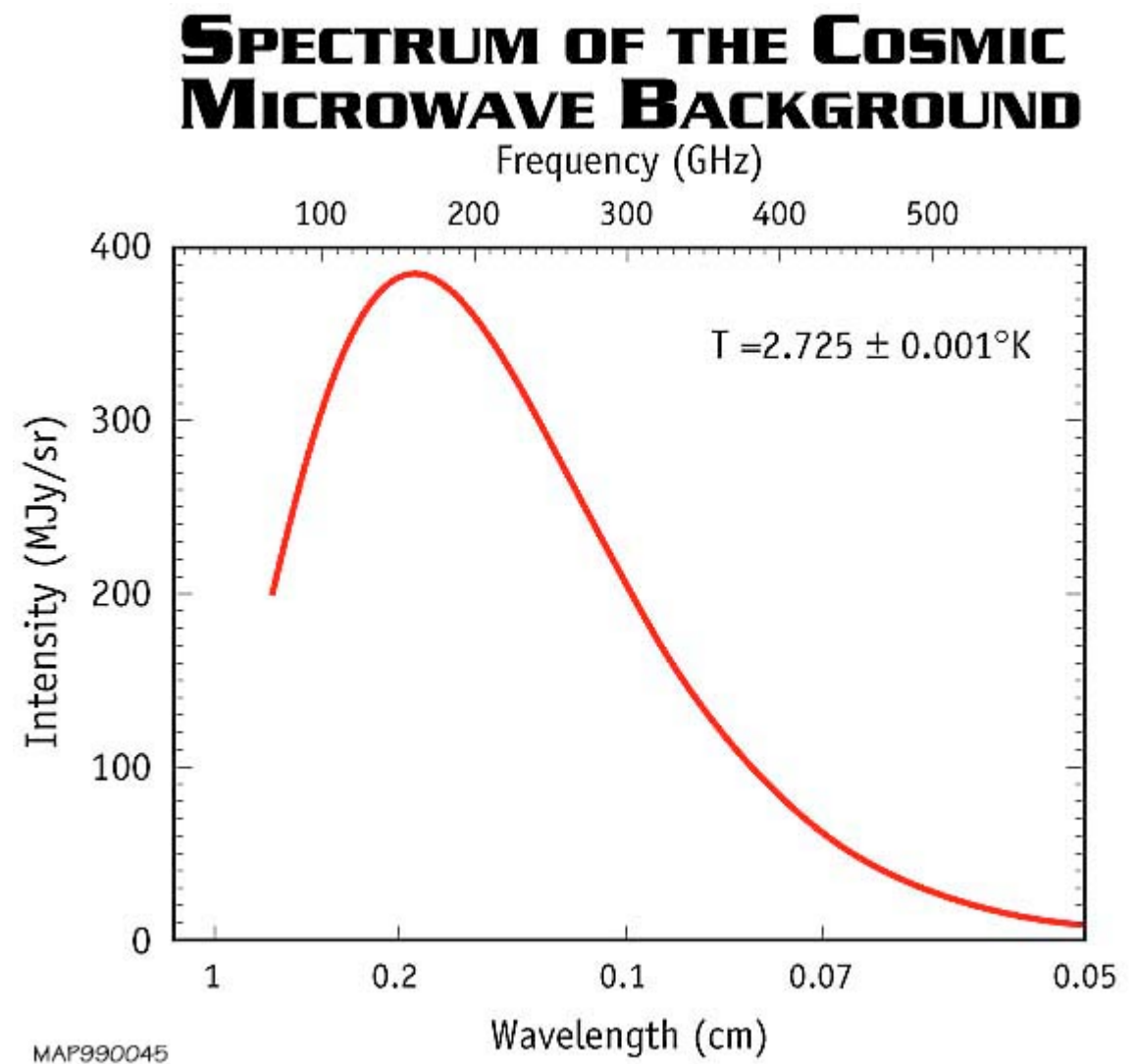
Ex: The peak wavelength of the CMB radiation at present is about 1 mm. In the past there was a time when the peak wavelength corresponded to blue light (400 nm). Estimate the scale factor ratio.

$$z = \frac{\Delta\lambda}{\lambda_0} = \frac{R}{R_0} - 1$$

$$\frac{R}{R_0} = \frac{\Delta\lambda}{\lambda_0} + 1 = \frac{10^{-3} - 400 \times 10^{-9}}{400 \times 10^{-9}} + 1 = 2500$$

# CMBR

- Ex: show that 2.7K blackbody radiation corresponds to microwave EM radiation



$$\lambda_{\text{max}} T = 2.9 \times 10^{-3} \text{ mK}$$

$$\lambda_{\text{max}} = \frac{2.9 \times 10^{-3} \text{ mK}}{2.725 \text{ K}} = 1.1 \text{ mm}$$

# Questions

- Work on option D questions from my website

